

# LISA on table, an electro-optical simulator for space based gravitational observatories

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## *Outline*

1. Gravitational waves

2. Gravitational waves detectors

3. LISA

4. LISA on table

# *Outline*

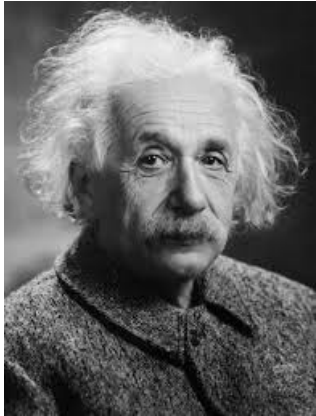
1. Gravitational waves

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## Gravitational waves



Predicted by Albert Einstein in his General Relativity Theory.

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu} \qquad R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$$
$$\square \bar{h}_{\mu\nu} = -\frac{16\pi G}{c^4}T_{\mu\nu}$$

Two different polarisations x and +, order of magnitude :  $10^{-21}$

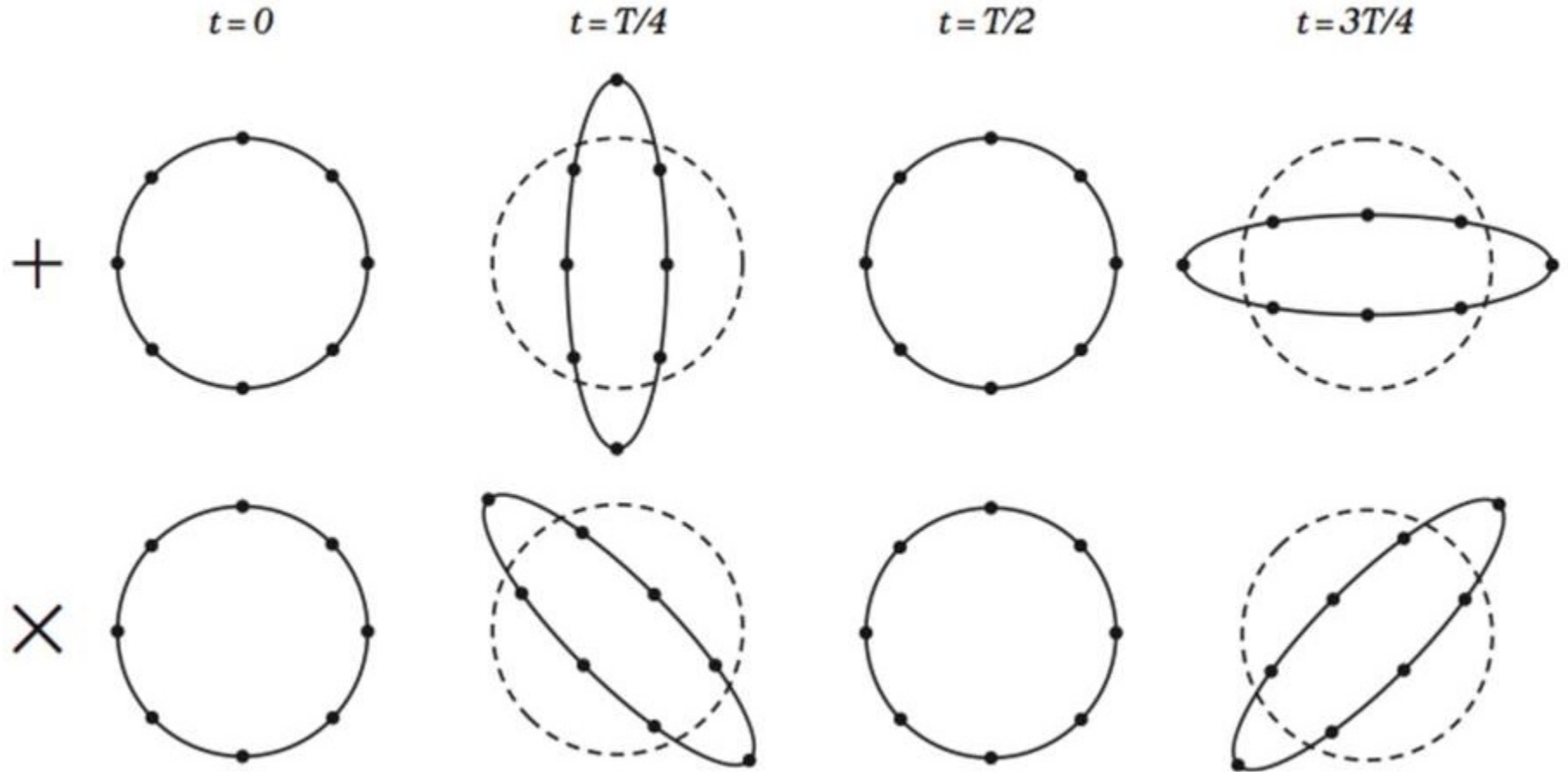
Gravitational waves are produced by massive aspheric accelerated systems and violent mass changes:

- Coalescing binaries (**black holes, neutron stars**),
- Magnetars,
- Supernovae...



## Gravitational waves

The effect of a gravitational wave on a set of free falling particles:



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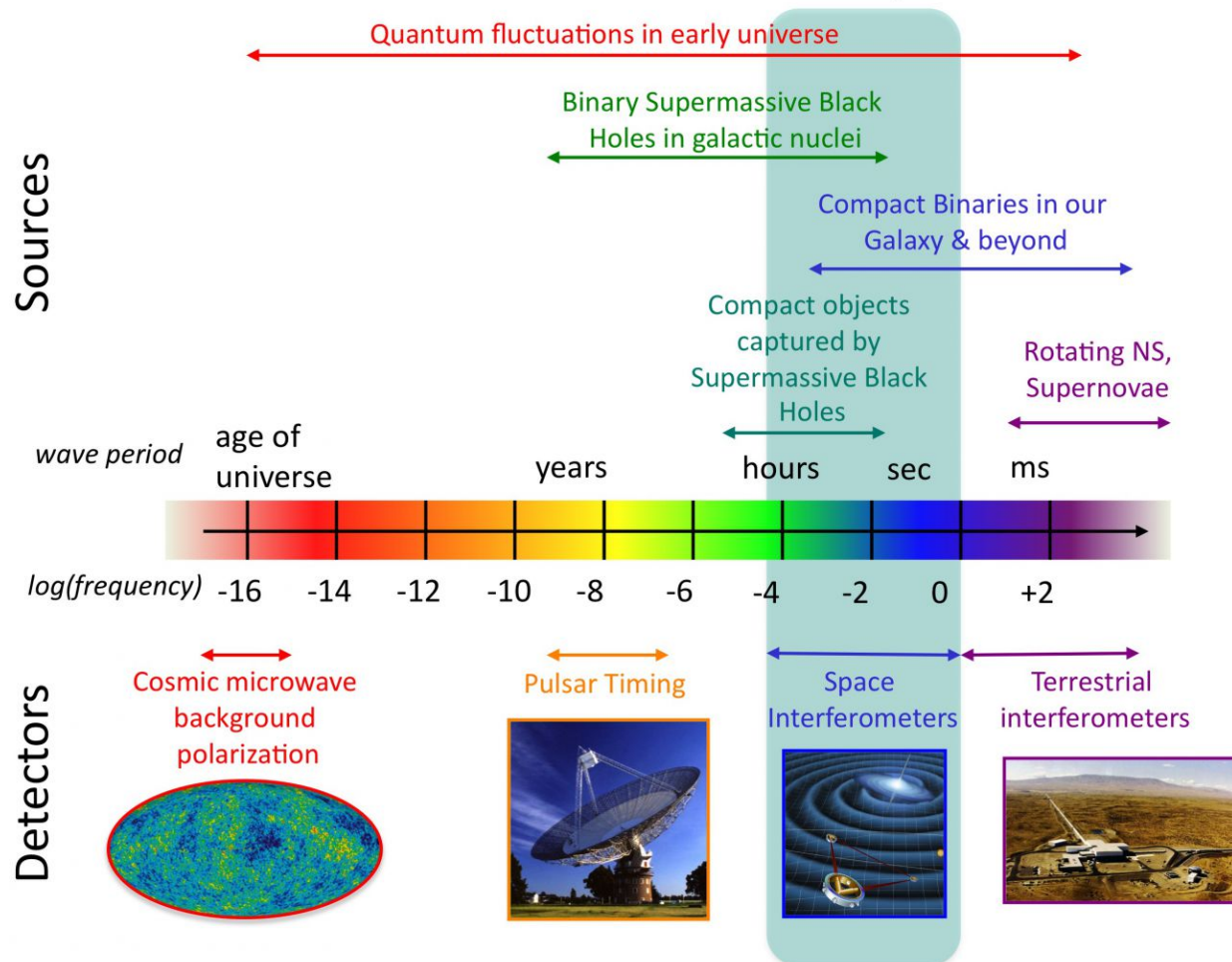
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# Gravitational waves detectors

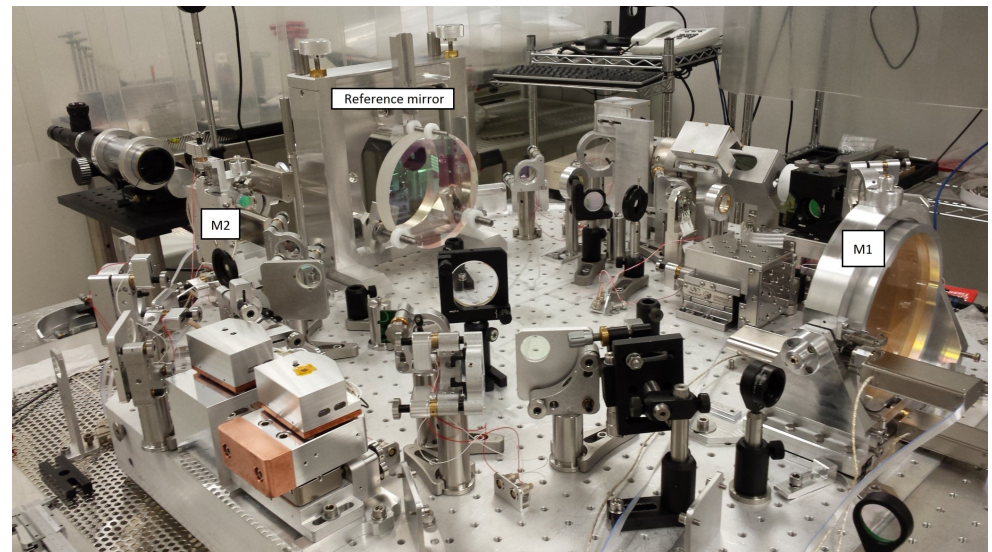
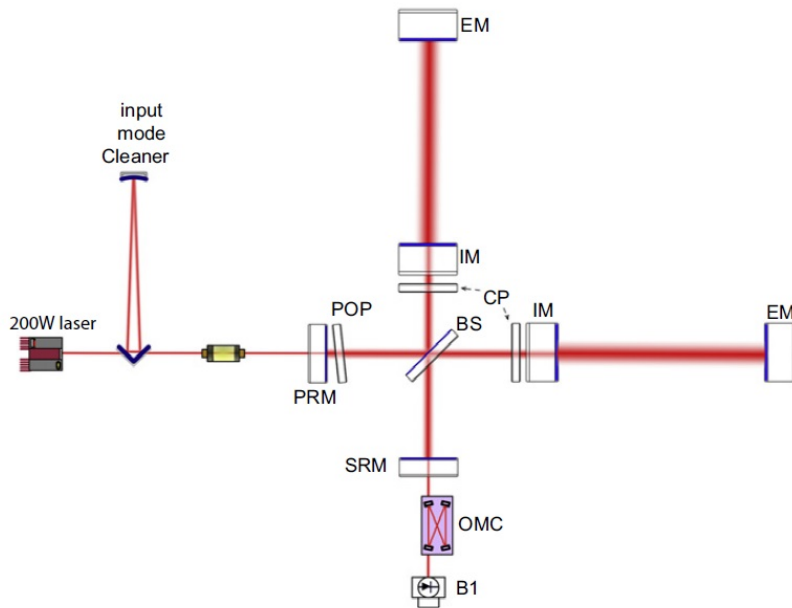
## The Gravitational Wave Spectrum



# Gravitational waves detectors

Ground detectors:

- LIGO (x2), USA
- Virgo, Italy
- KAGRA, Japan
- Einstein Telescope





## First Detection

- First detection by LIGO on september 14, 2015
- Binary black holes merger

Primary black hole mass

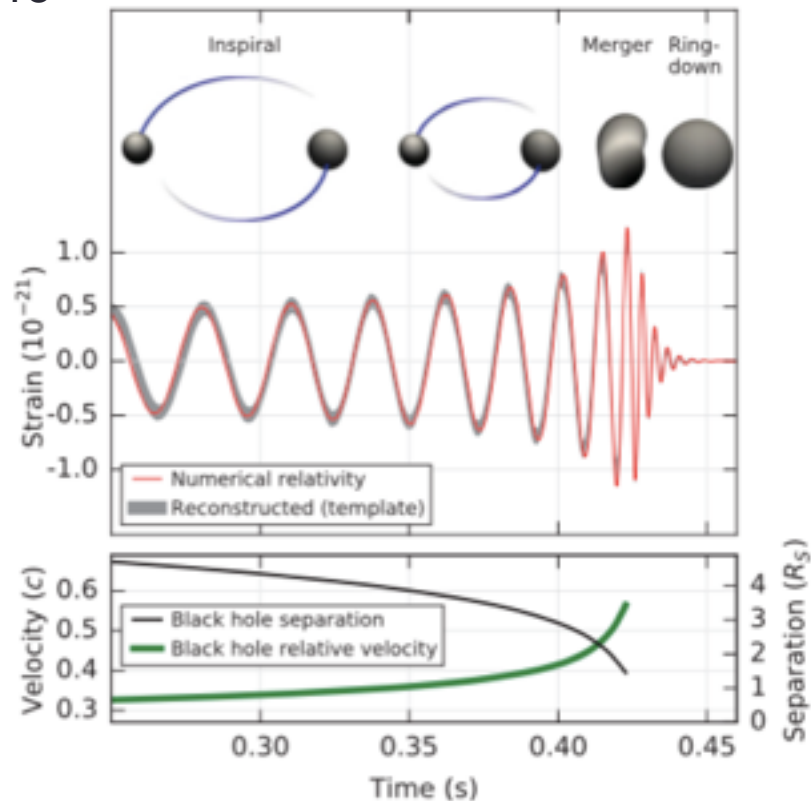
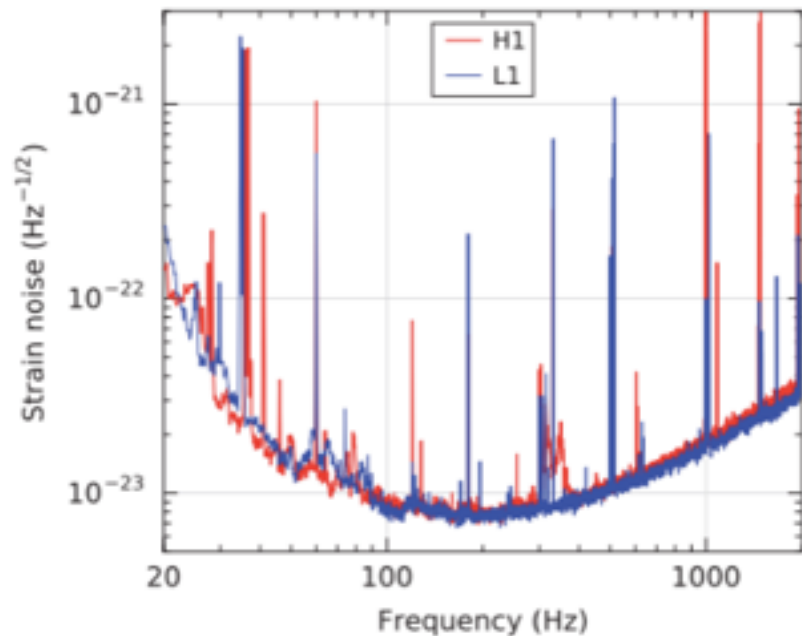
$$36^{+5}_{-4} M_{\odot}$$

Secondary black hole mass

$$29^{+4}_{-4} M_{\odot}$$

Final black hole mass

$$62^{+4}_{-4} M_{\odot}$$



B.P. Abot *et al.*, Observation of Gravitational Waves from a Binary Black Hole Merger, PRL 116, 061102 (2016)

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**3. LISA**

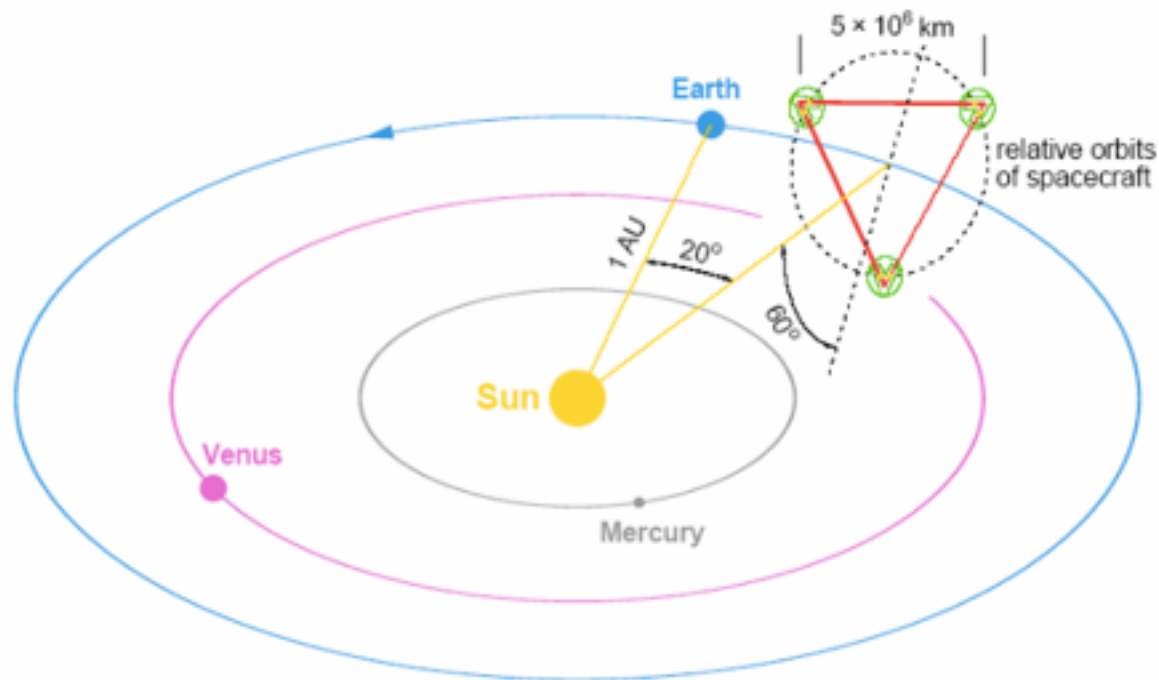
4. LISA on table

# LISA

Mentioned in the 90s, launch in 2034, ESA-(NASA) mission.

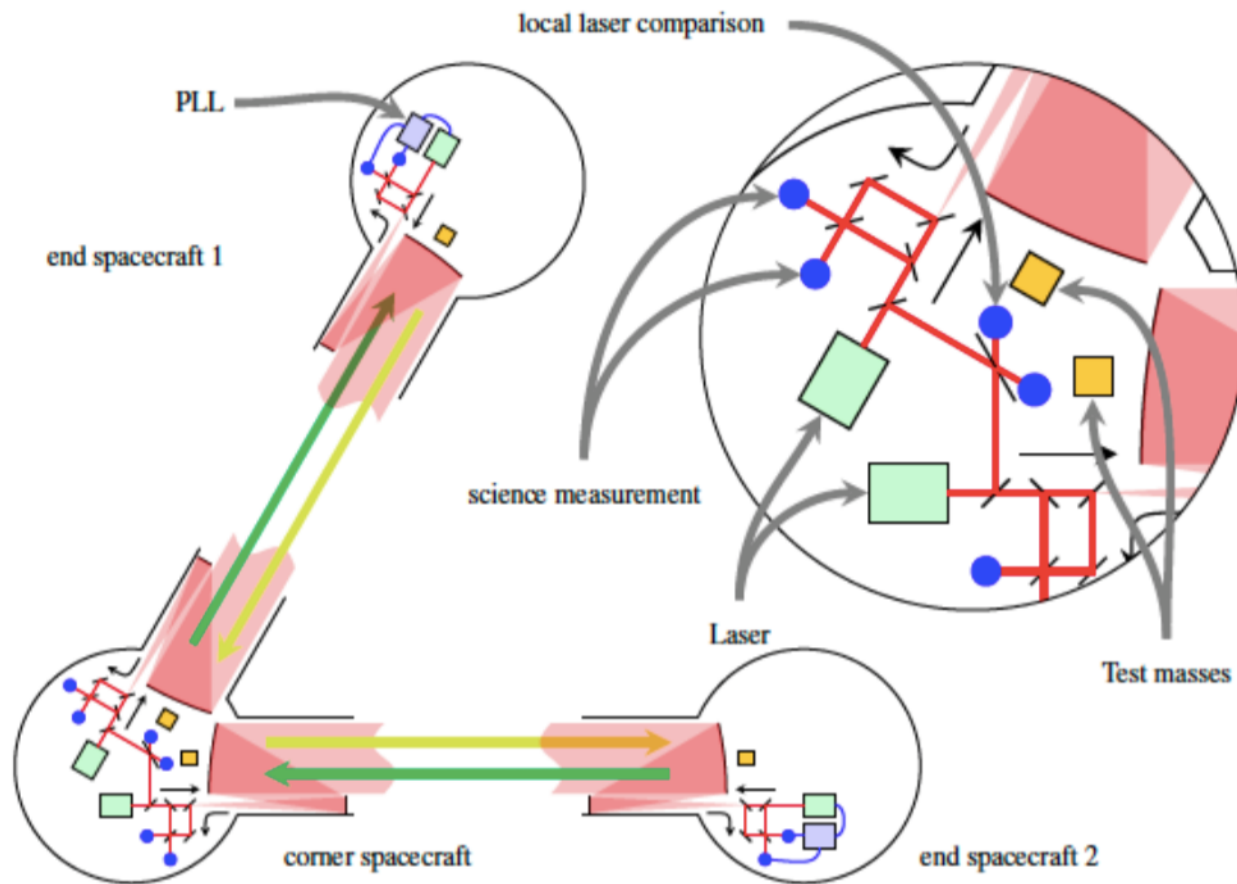
Three satellites separated by few millions km, forming an equilateral triangle.

Orbital configuration:



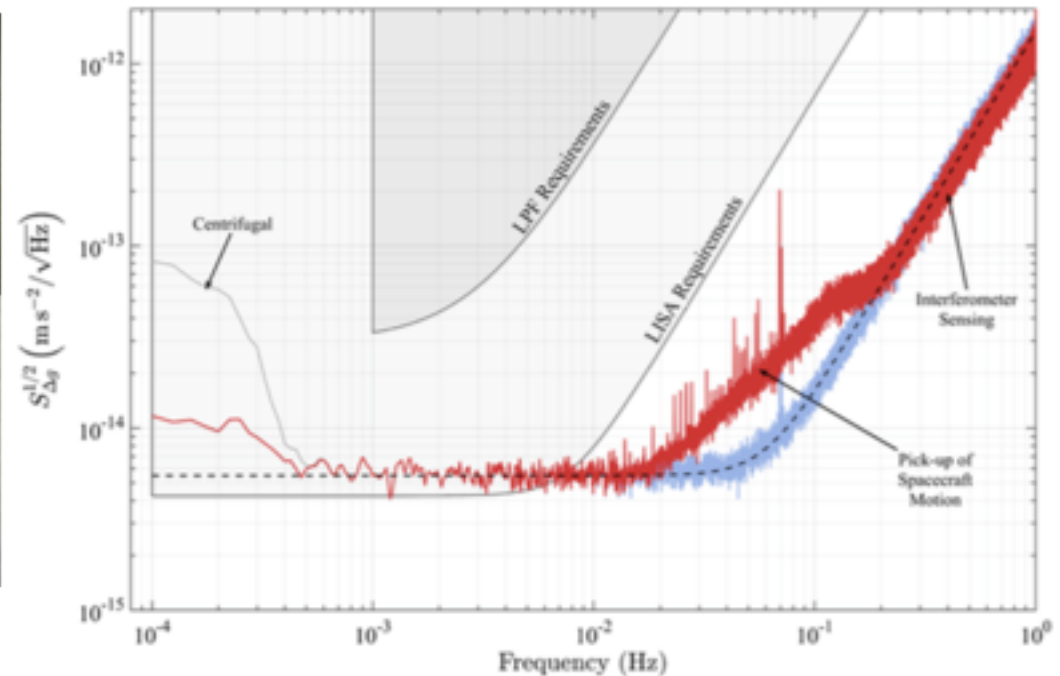
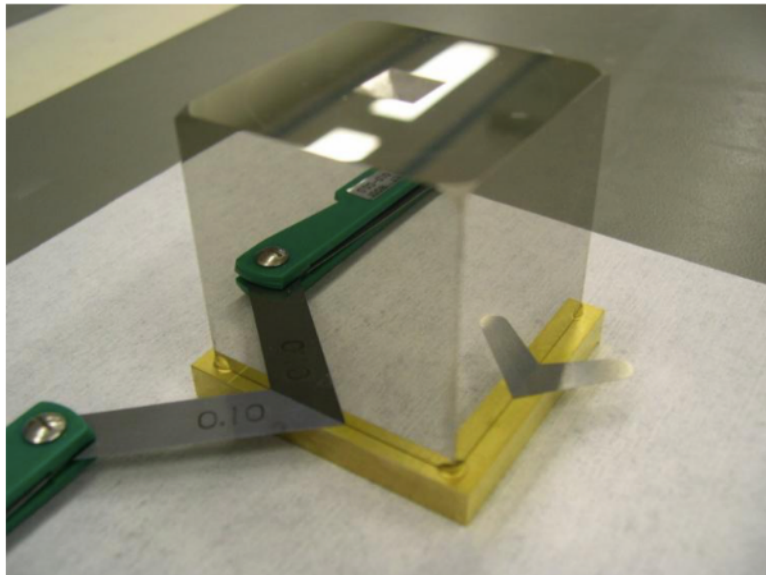
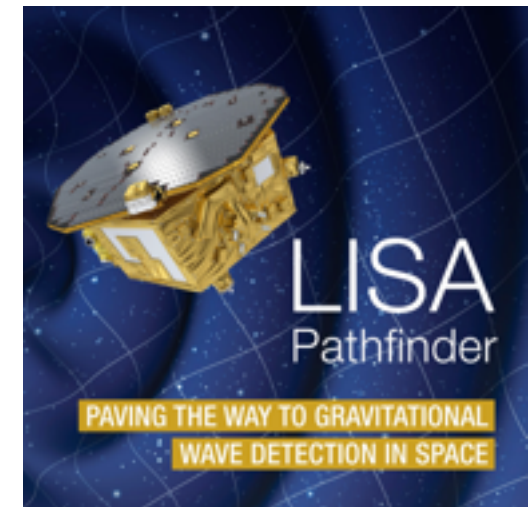
# LISA

Simplified scheme of the constellation:



## LISA Pathfinder

- Launched in 2016
- Demonstrates the feasibility of the drag-free principle
- Residual acceleration of test masses beyond LISA requirements



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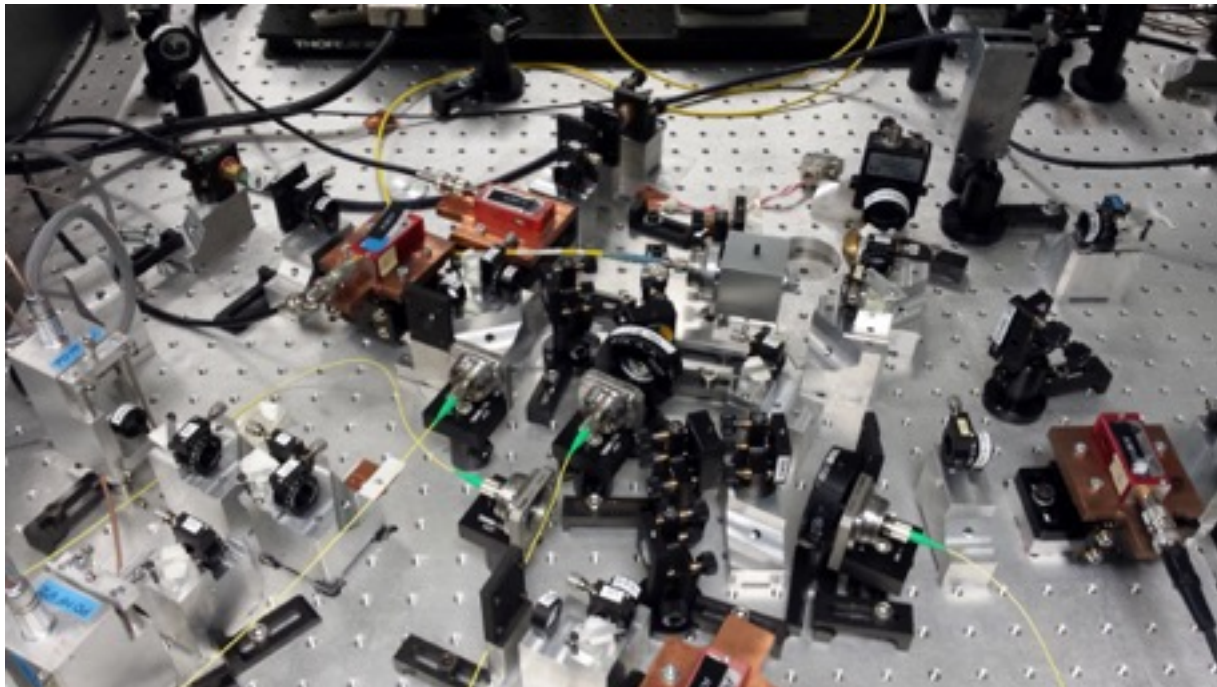
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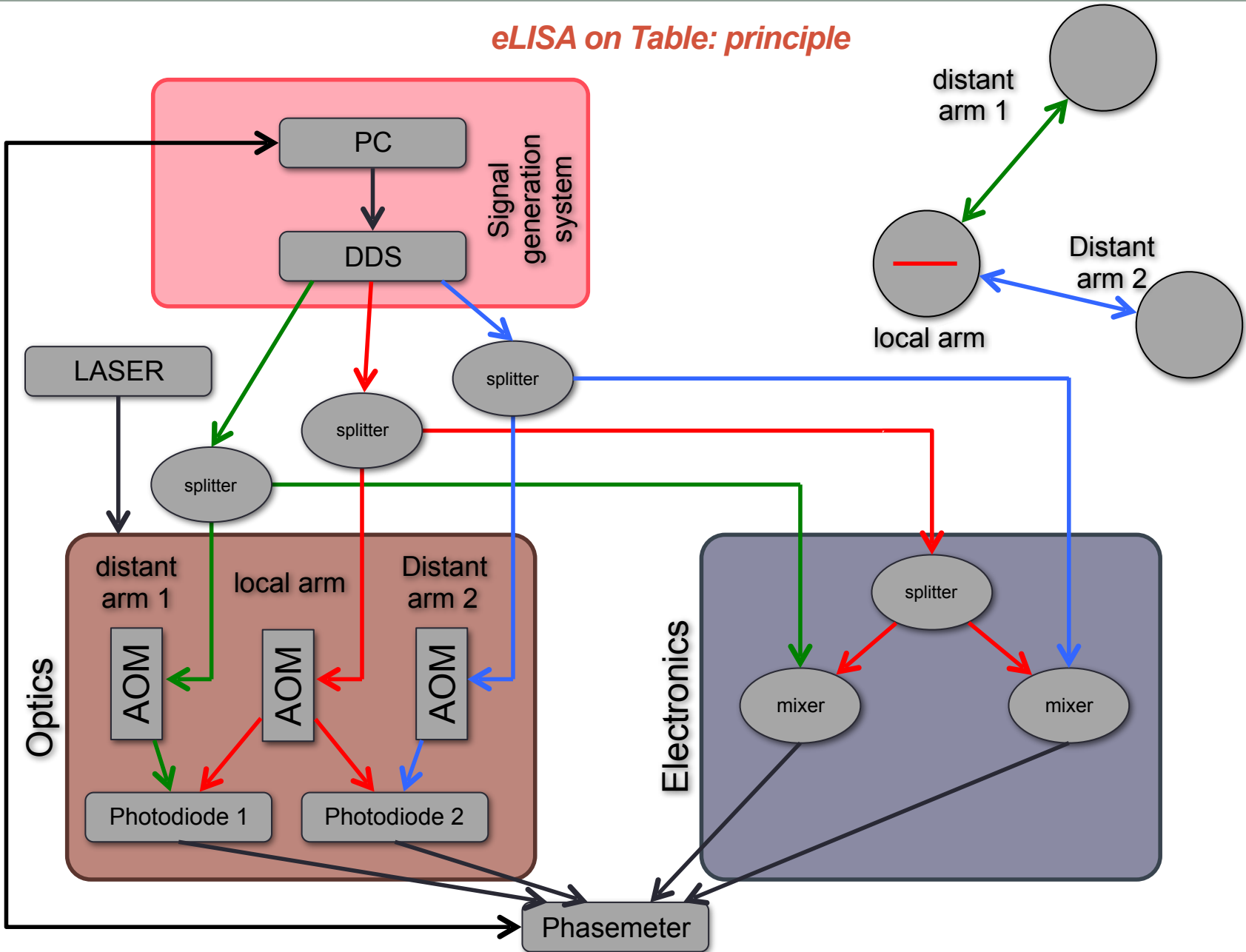
4. LISA on table

## *LISA on Table*

- Optical and electronic simulator of LISA.
- Objectives: to test the noise reduction techniques experimentally, to test instruments (photodiodes, phasemeter, ...) in a representative acquisition chain.

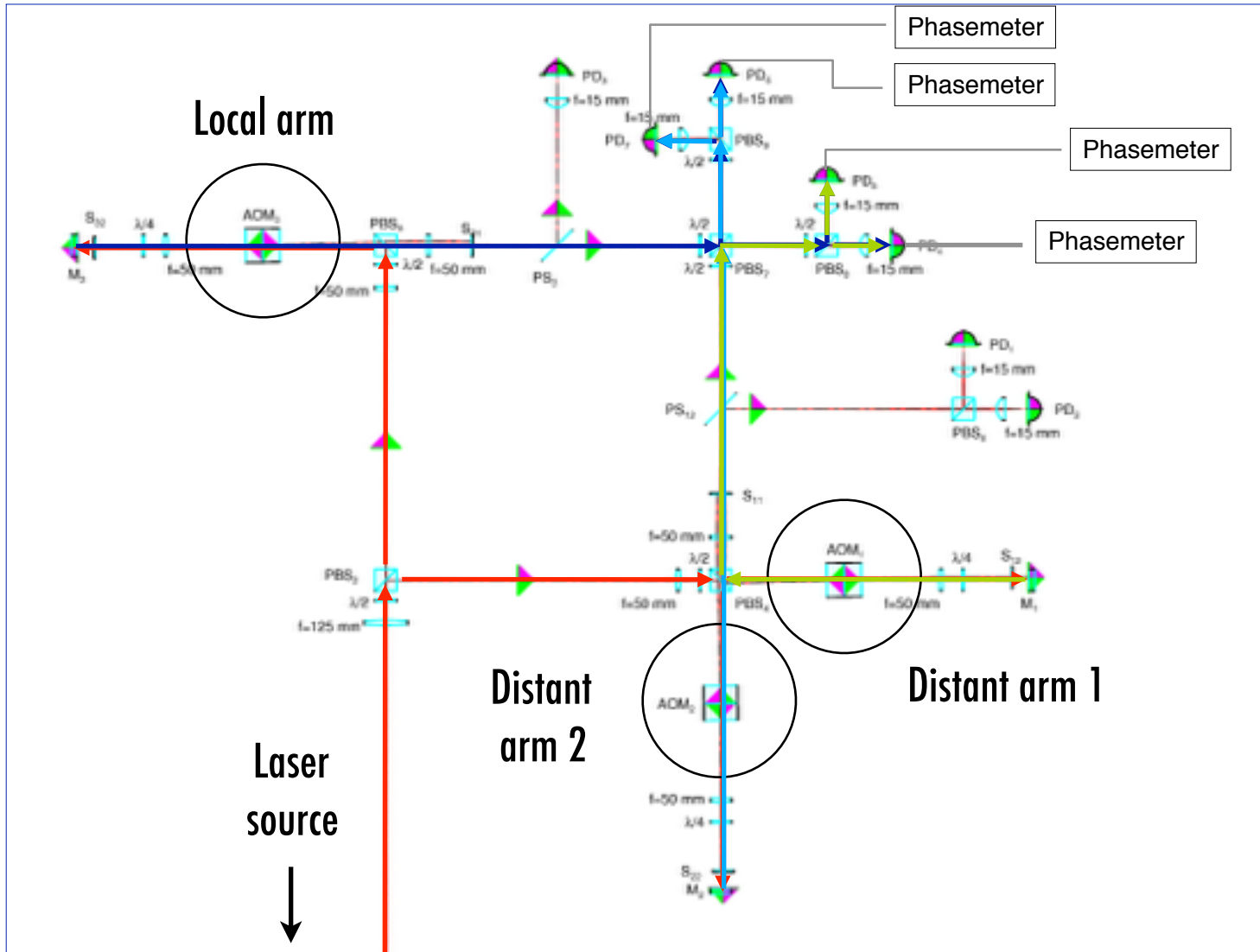


*eLISA on Table: principle*

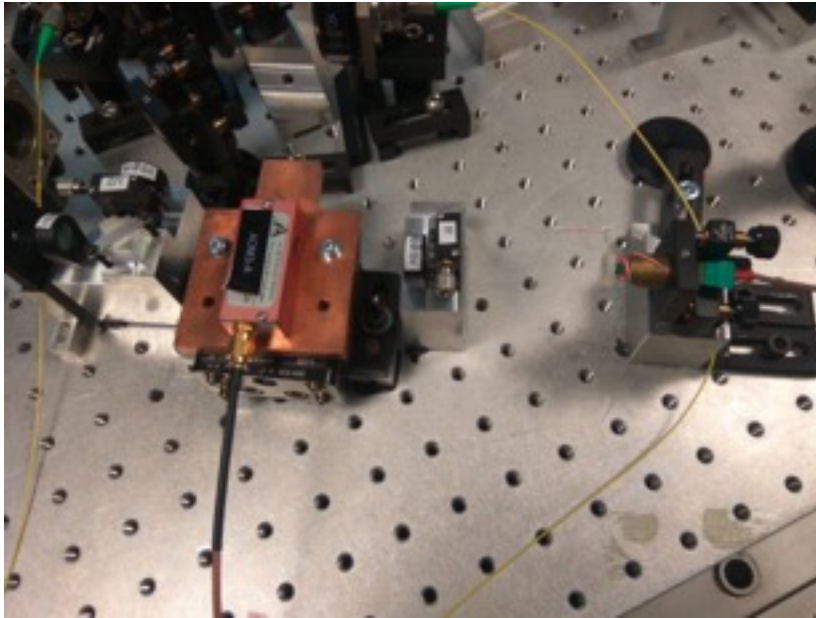




## eLISA on Table: optical layout

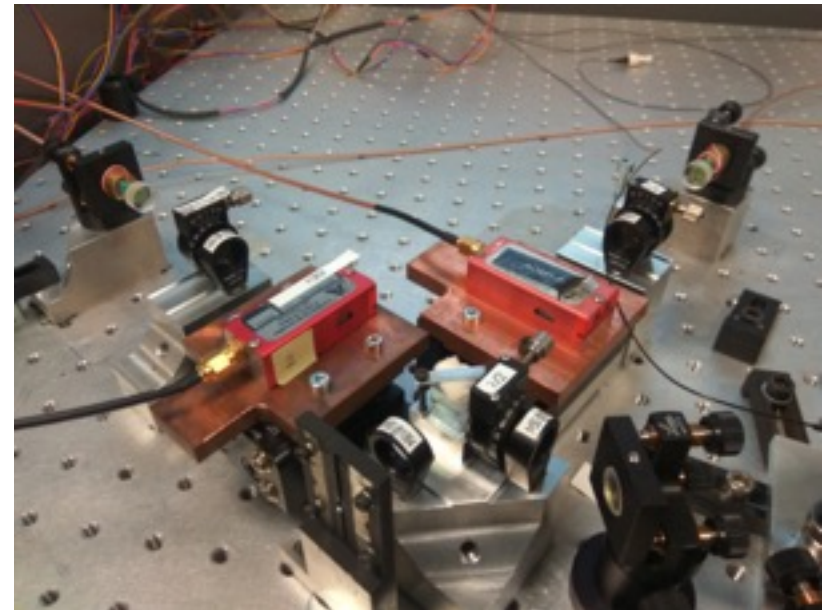
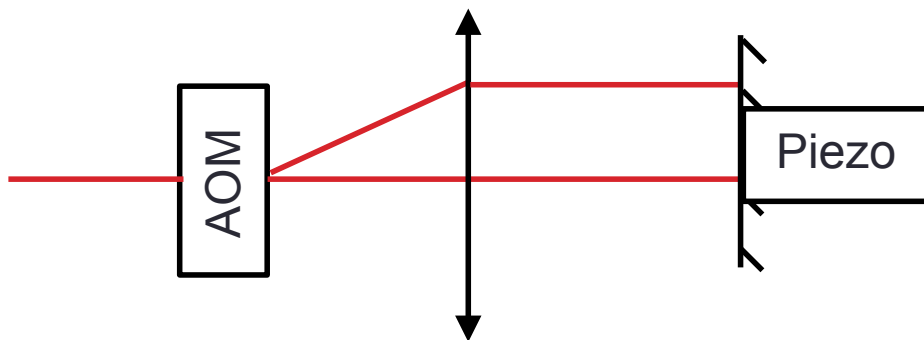


## *eLISA on Table*



Bragg's cell in cat's eye configuration:

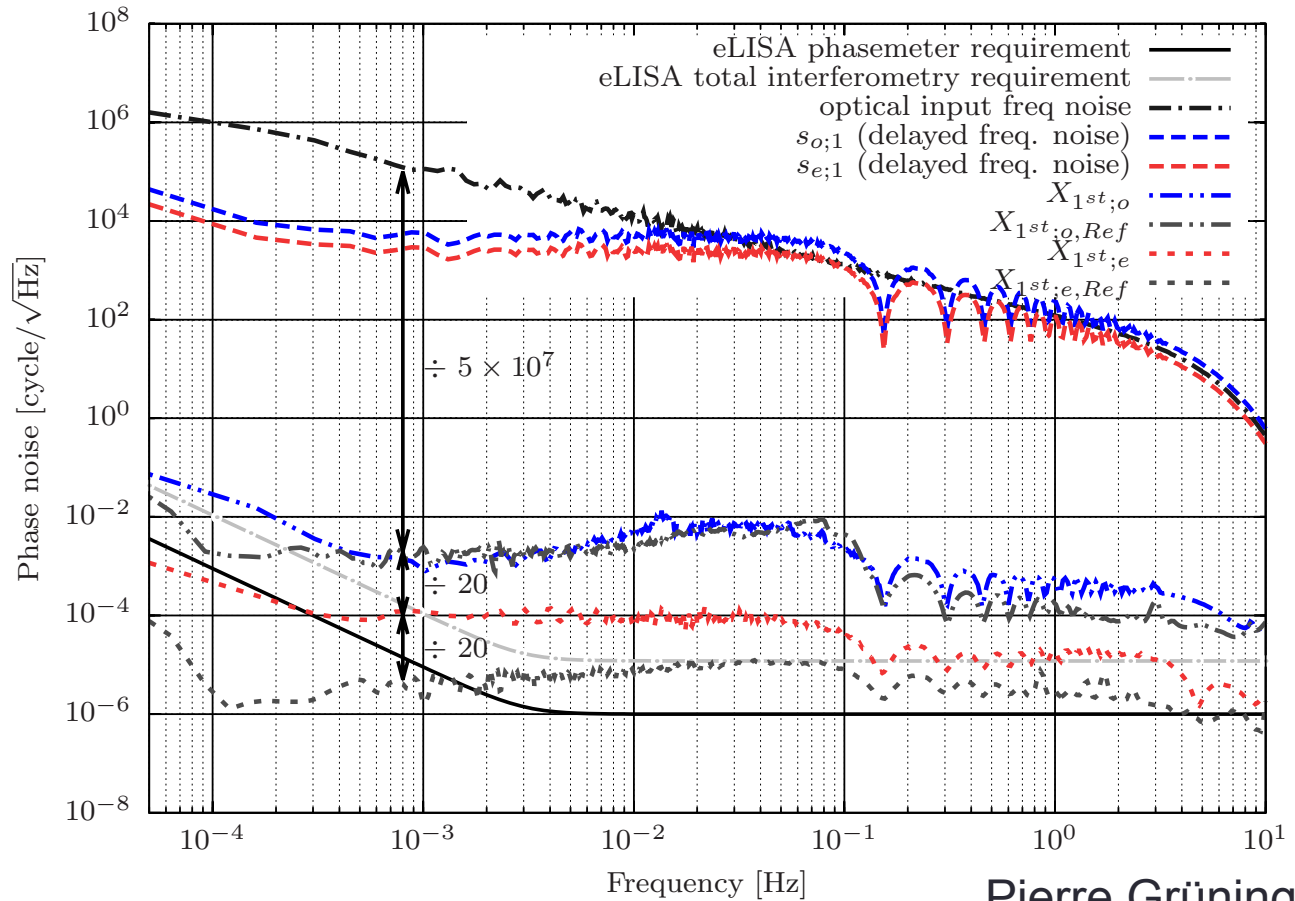
The diffracted beam is modulated in frequency



## eLISA on Table

results for both interferometers in the following configuration:

- TDI 1st generation,
- static, uneven arms,
- white noise.



Pierre Grüning

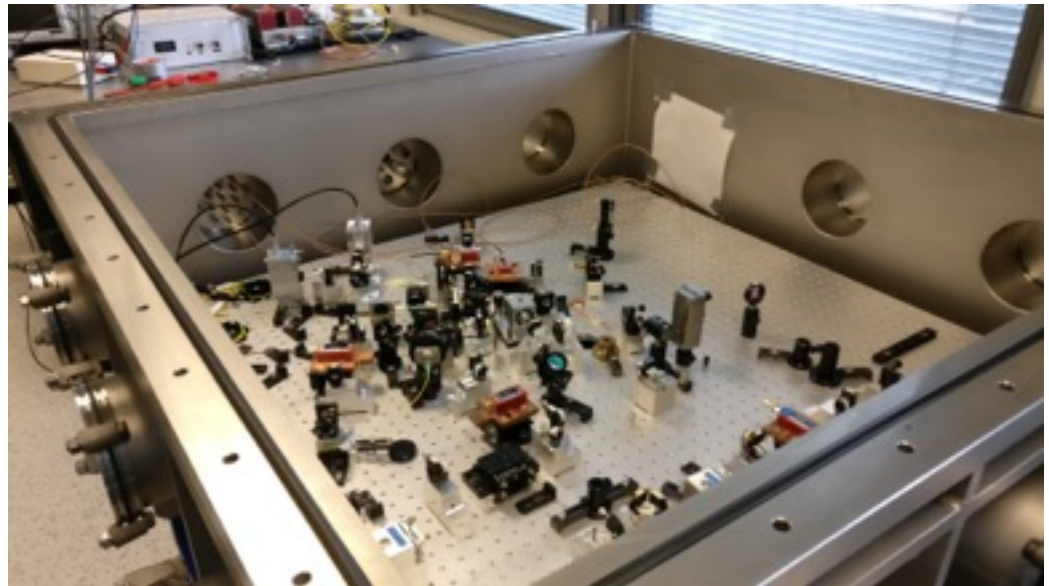
## *LISA on Table*

### **Optical interferometer:**

Only limited by the system,  
which means TDI 1st  
generation works

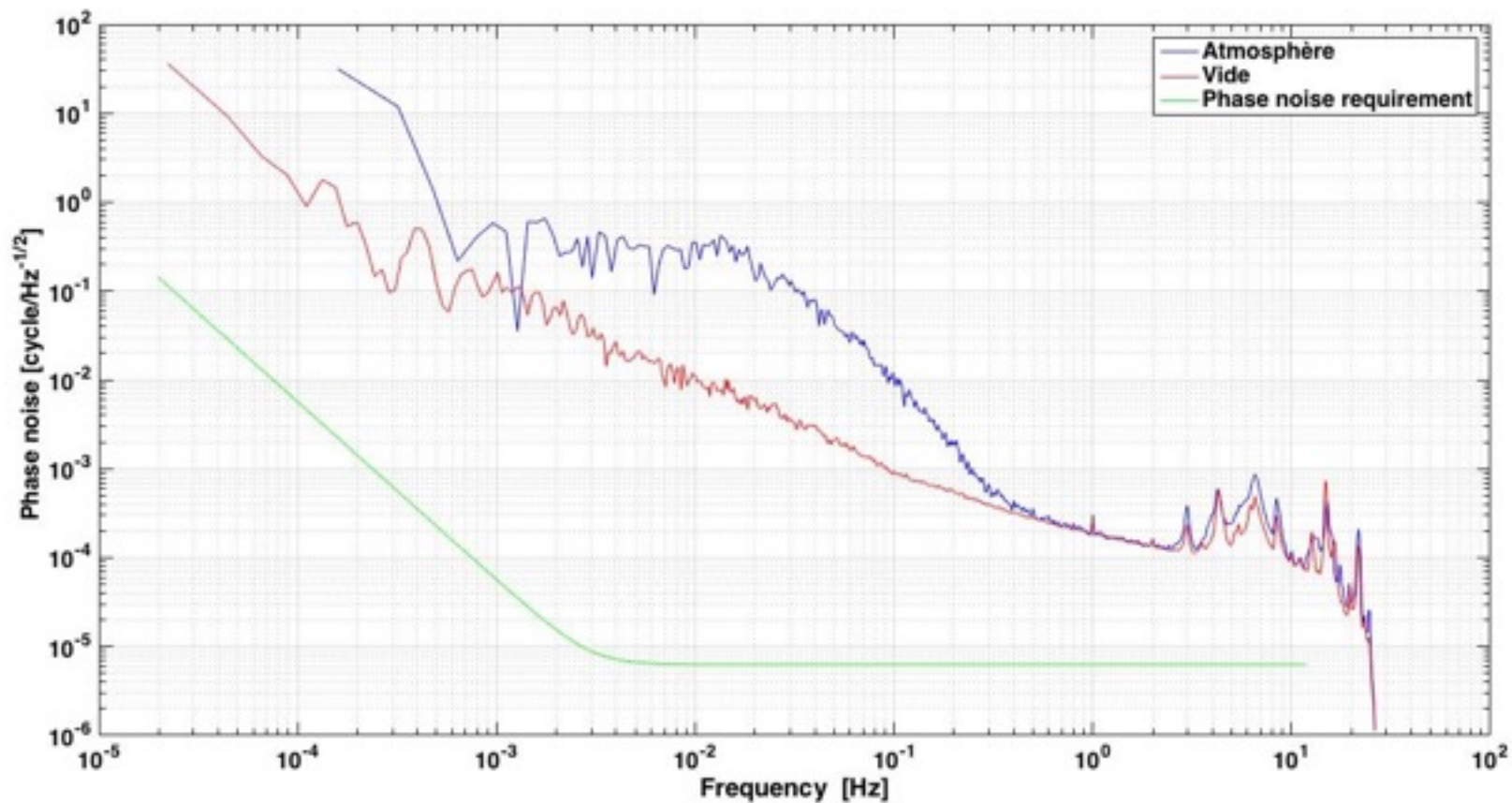
LOT is now in a vacuum  
chamber to further reduce  
optical noise

This requires a lot of  
experimental work



## LISA on Table

### Vacuum operation validation

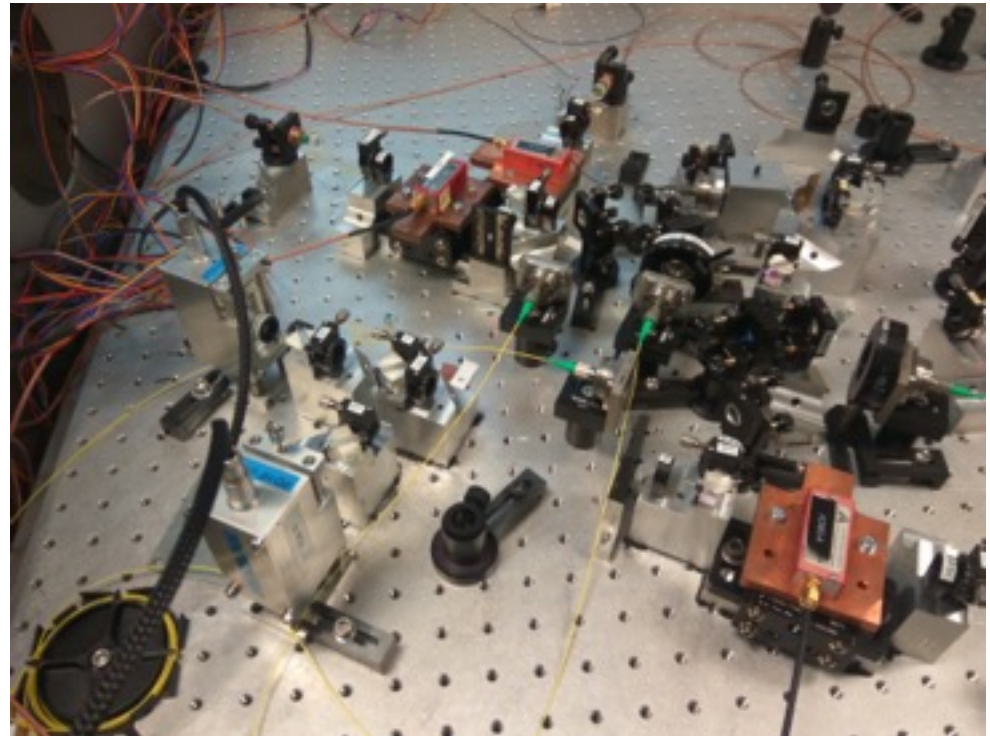




## *LISA on Table*

On-going work:

- Active compensation optimisation
- Optical system noise investigation
- Doppler effect simulations



***Thank you!***